The University of Alabama at Birmingham (UAB) **Department of Physics**

PH 462/562 - Classical Mechanics II - Spring 2006

Central Force Problem: Summary & Highlights

General Form of a Central Force: $\vec{F} = F(r)\hat{e}_r$

(1) $\nabla \times \vec{F} = 0 \Rightarrow \vec{F}$ is conservative $\Rightarrow E = const$

(2)
$$\frac{d\vec{L}}{dt} = \frac{d}{dt} (\vec{r} \times \vec{p}) = 0 \Rightarrow \vec{L} = const (motion takes place in plane)$$

Hence, polar coordinate system very convenient: r(t) = ?; $\theta(t) = ?$

$$L = mr^2 \dot{\theta} = const$$

Complementary Approaches to Solve the Central Force Problem

(1) Equations of Motion in Polar Coordinates

$$m\ddot{r} - mr\dot{\theta}^2 = F(r)$$
$$mr\ddot{\theta} + 2m\dot{r}\dot{\theta} = 0$$

(2) Conservation of Energy

$$E = T + V = \frac{1}{2}mv^{2} + V(r) \Rightarrow \sqrt{\frac{2}{m}}t = \int_{r_{0}}^{r} \frac{dr}{\sqrt{E - V(r) - \frac{L^{2}}{2mr^{2}}}} \Rightarrow$$
$$\Rightarrow r(t); \quad \theta(t) \quad \left[from \ \dot{\theta} = \frac{L}{mr^{2}}\right]$$

(3) Analysis based on Effective Potential Energy

$$U_{eff}(r) = V(r) + \frac{L^2}{2mr^2}$$

(4) Differential Equation for Orbit (i.e., trajectory in real space)

$$\frac{d^2 u}{d\theta^2} = -u - \frac{m}{L^2 u^2} F\left(\frac{1}{u}\right) \implies u(\theta) = \frac{1}{r(\theta)}$$

Assignment # 9 Due: Friday, April 28, 2006

- 1. A particle of mass *m* moves in three dimensions with a potential energy $V(r) = \frac{1}{2}kr^2$, where *r* is the radial coordinate of the spherical coordinate system, and *k* is a positive constant.
 - a. Make a sketch showing V(r), the centrifugal potential energy $\frac{L^2}{2mr^2}$, and the effective potential energy $U_{eff}(r)$ (*Treat the angular momentum L as a known, fixed constant, that can be determined from the initial conditions*).
 - b. Find the equilibrium distance r_0 (i.e., the distance at which the particle can circle the origin with constant radius).
 - c. By making a Taylor expansion of $U_{eff}(r)$ about the equilibrium point r_0 and neglecting all terms in $(r r_0)^3$ and higher, find the frequency of small oscillations about the circular orbit if the particles are disturbed a little from the separation r_0 .
 - d. Set up and solve the differential equation of the orbit and find the general form of the trajectories in three-dimensional space.